Modeling the bimodal orientation of filamentary molecular clouds

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Introduction: molecular clouds form filaments



Example: IC 5146 / Barnard 168

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Observation: bimodal cloud orientations



H-b. Li et al. MNRAS 436, 3707 (2013)

- study of 13 filamentary molecular clouds in the Gould Belt
- clouds are aligned either parallel or perpendicular to the magnetic field of the local inter-cloud medium
- strong significance: probability for this correlation to occur from random orientations < 0.6%

Data Sources:

- filamentary structures determined from dust extinction maps (Dobashi 2011)
- ICM B-field direction obtained from optical stellar polarimetry data (Heiles 2000)

Interpretation: B-fields are dynamically important

Type I $P_{turb} > P_{grav}$

B-field channels anisotropic sub-Alfvénic turbulence







B-field channels gravitational contraction



Computational Setup

- Code: ZEUS-MP + modifications
- Simulation volume: (300pc)³
- Grid: 128³ ... 512³, periodic
- Turbulence [Stone et al. ApJ 508, L99 (1998)]

$$\delta \mathbf{v}_{k}^{2} \propto k^{6} \exp(-8k/k_{0})$$
$$\nabla \cdot \delta \mathbf{v} = 0$$
$$\int \rho \delta \mathbf{v} d^{3} x = 0$$
$$dE/dt = \text{const}$$

- isothermal
- uniform initial density, B-field

What conditions lead to bimodal field-cloud alignment?



Preliminary Results



Preliminary Results



Ζ

- low mass
- moderate turbulence
- strong B-field



- many filaments at various scales
- B-field unaffected
- filaments mostly orthogonal to B-field

Outlook

- automatic filament identification & correlation (WIP)
- simulate single filaments
- find clumps inside filaments: classify as sub-critical / super-critical
- correlation between filament alignment and SFE?

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